

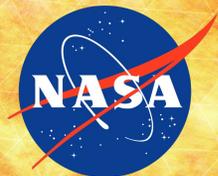
# The Origin of Weakened Magnetic Braking

**Travis Metcalfe**

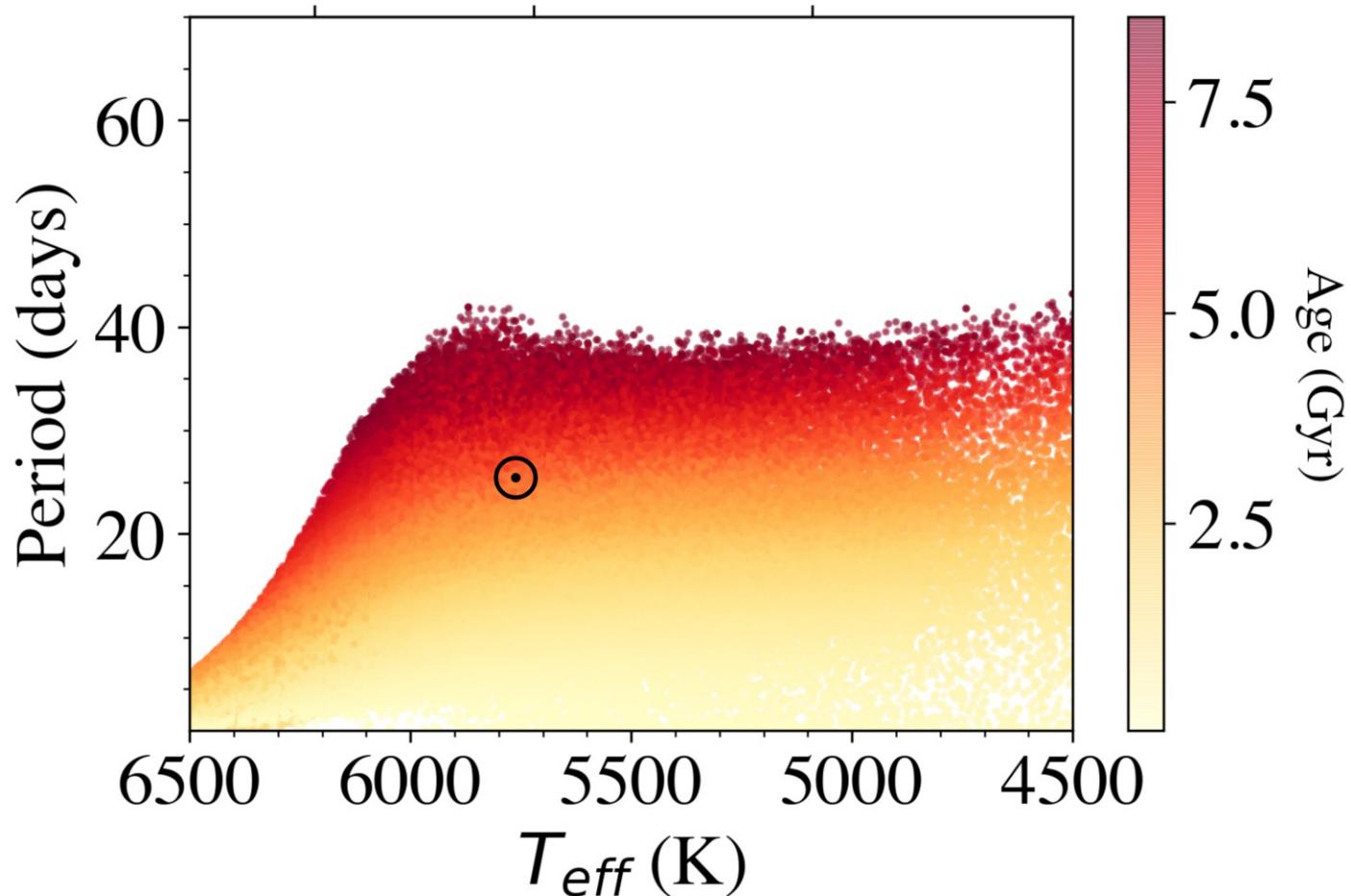
*White Dwarf Research Corp.*

[2022: ApJL 933, L17]

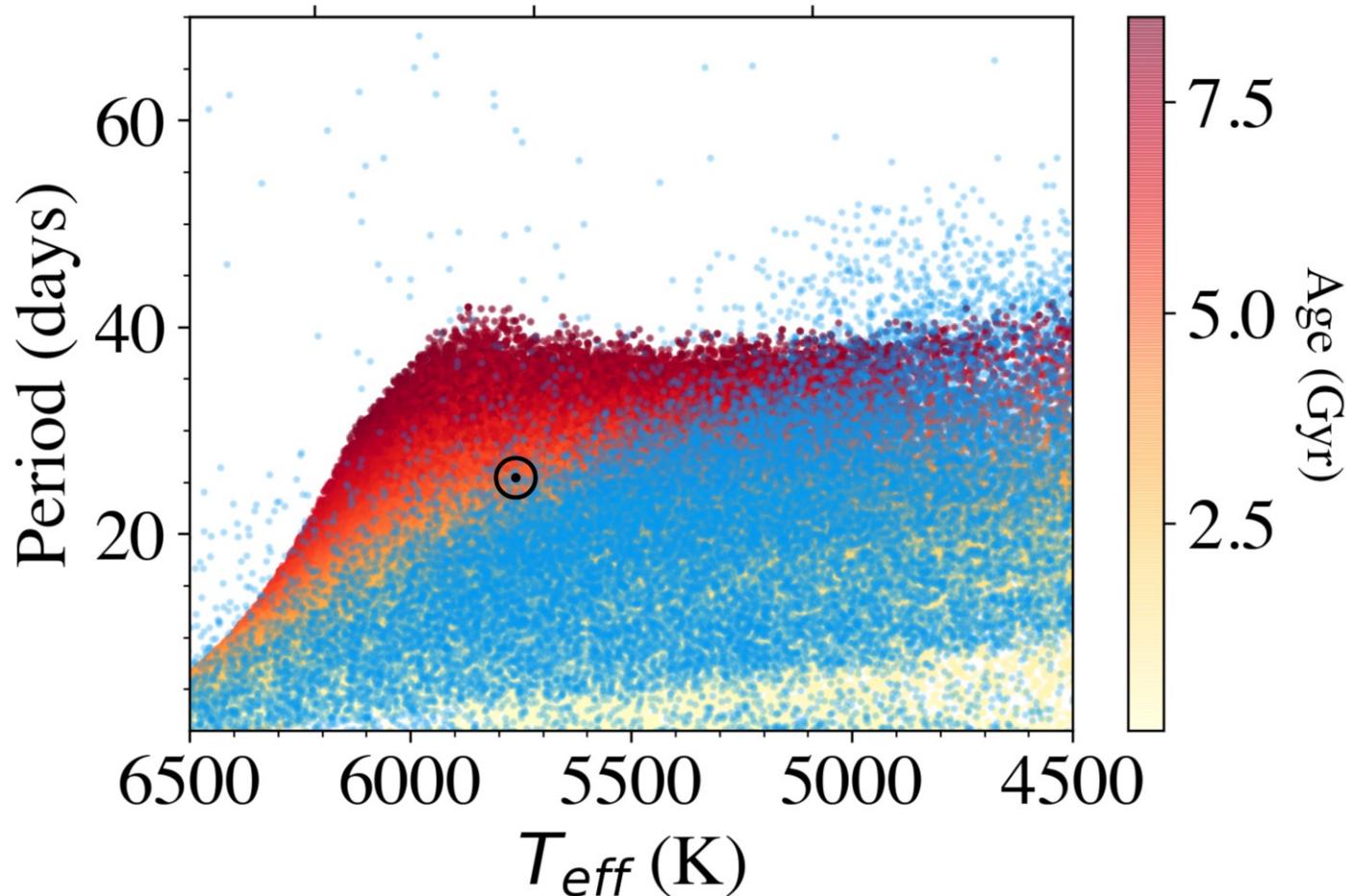
Collaborators: Adam Finley, Oleg Kochukhov, Victor See, Thomas Ayres, Keivan Stassun, Jennifer van Saders, Catherine Clark, Diego Godoy-Rivera, Ilya Ilyin, Marc Pinsonneault, Klaus Strassmeier, Pascal Petit



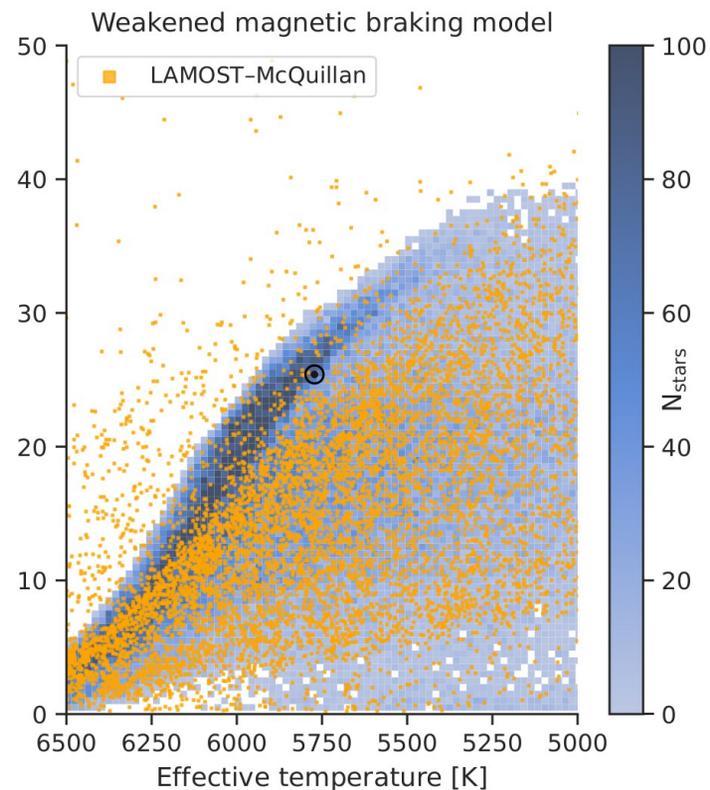
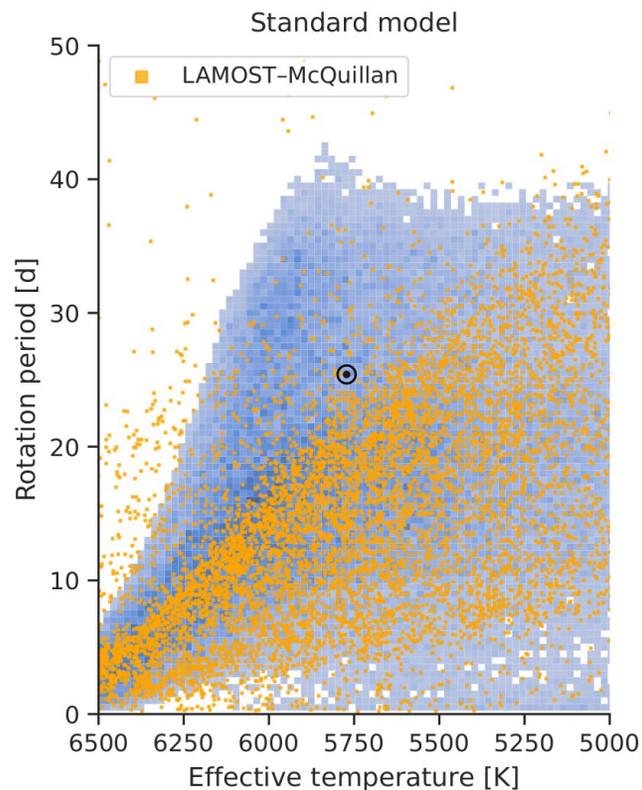
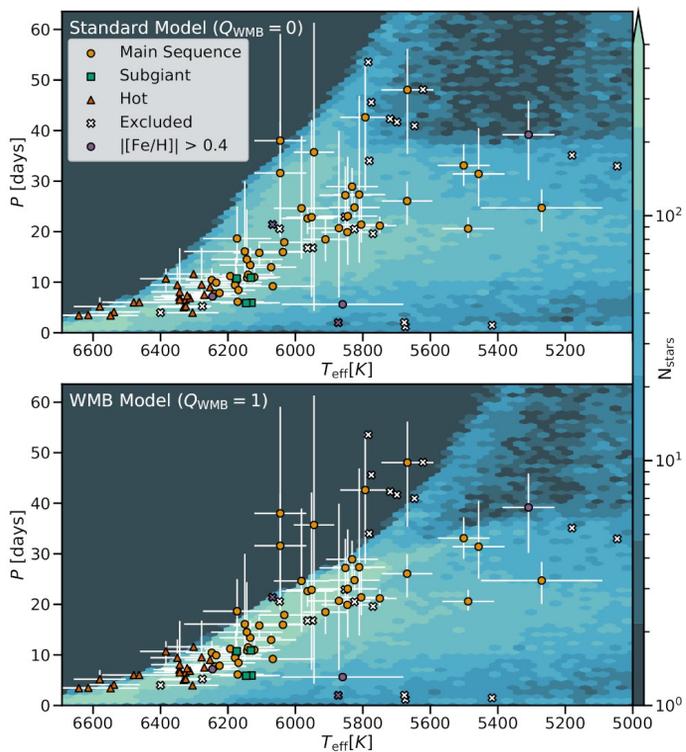
# Weakened magnetic braking



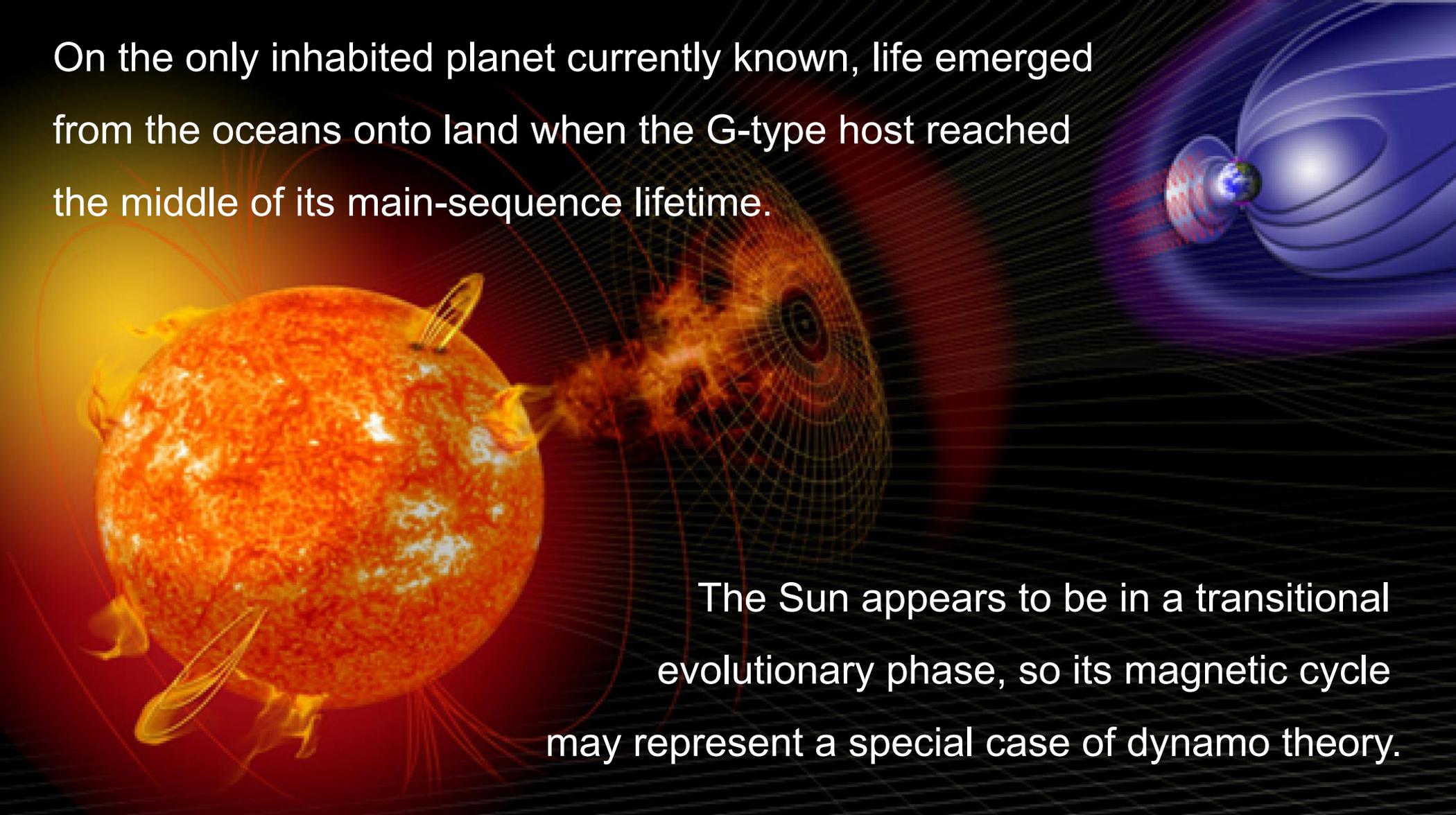
# Weakened magnetic braking



- **No detection bias:** asteroseismic rotation rates show similar distribution
- **Pile-up confirmed:** precise  $T_{\text{eff}}$  shows range of ages at long-period edge



On the only inhabited planet currently known, life emerged from the oceans onto land when the G-type host reached the middle of its main-sequence lifetime.

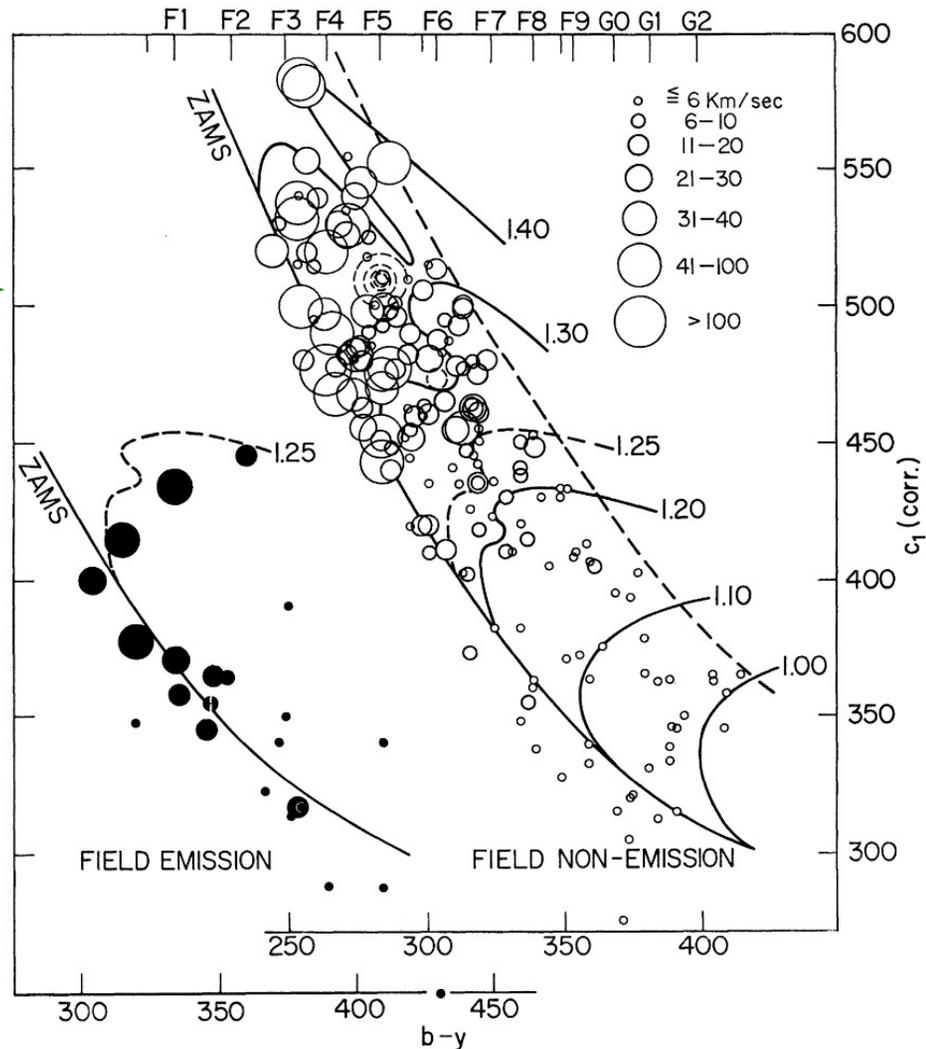
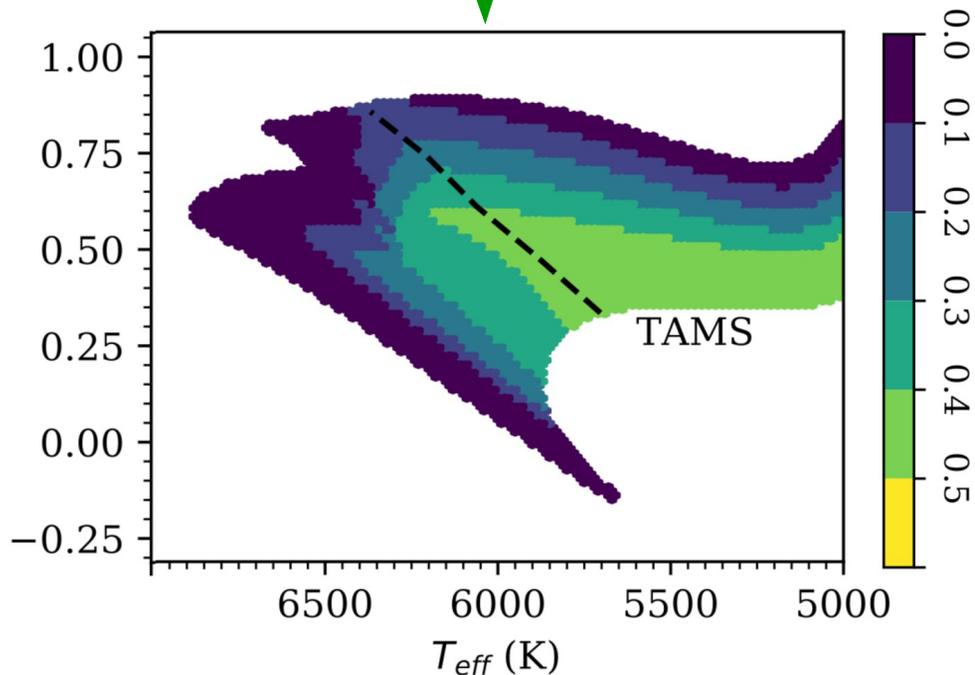


The Sun appears to be in a transitional evolutionary phase, so its magnetic cycle may represent a special case of dynamo theory.

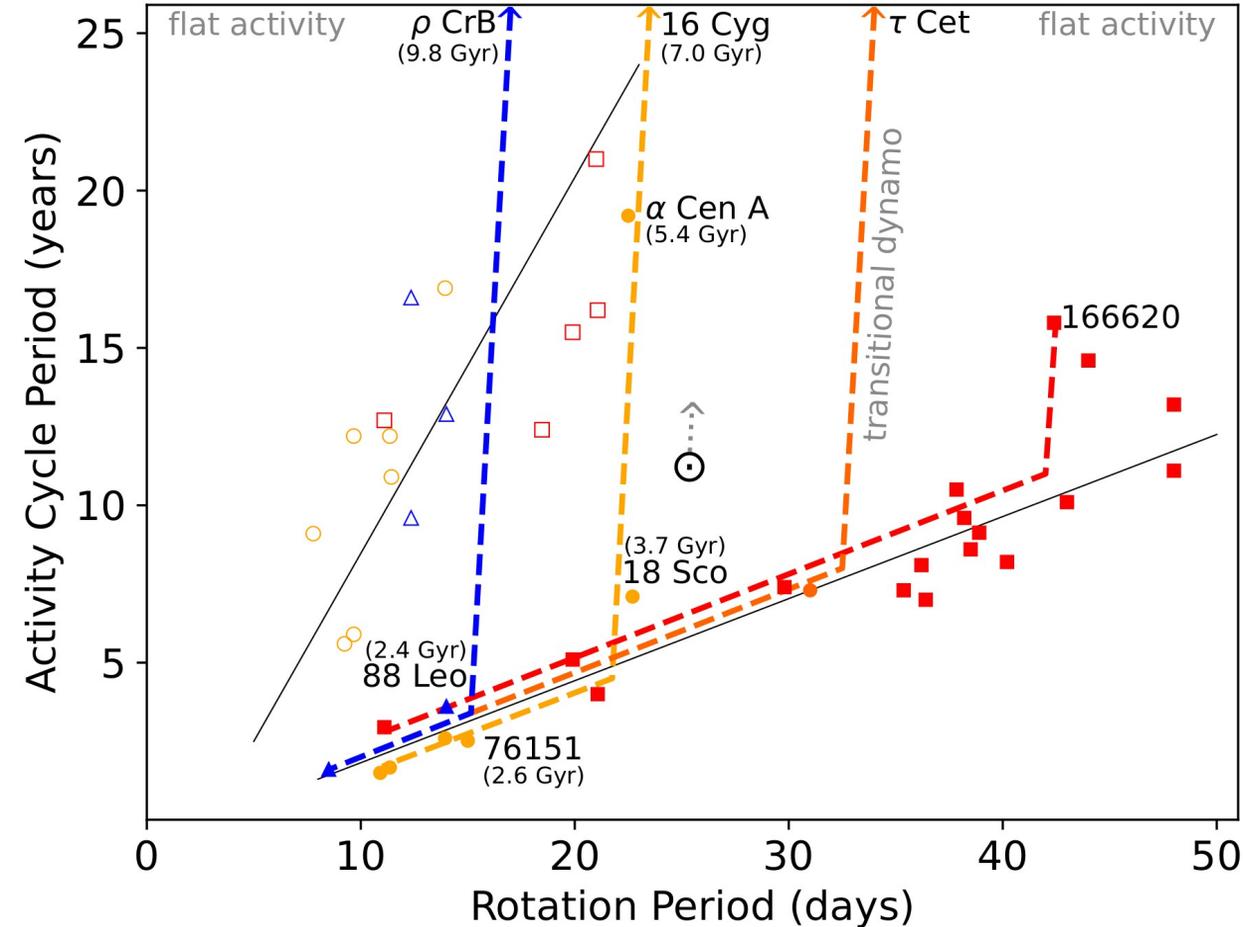
To sustain a magnetic dynamo, the Rossby number cannot be too large!

$$Ro \equiv P_{\text{rot}} / \tau_c$$

too long
too short

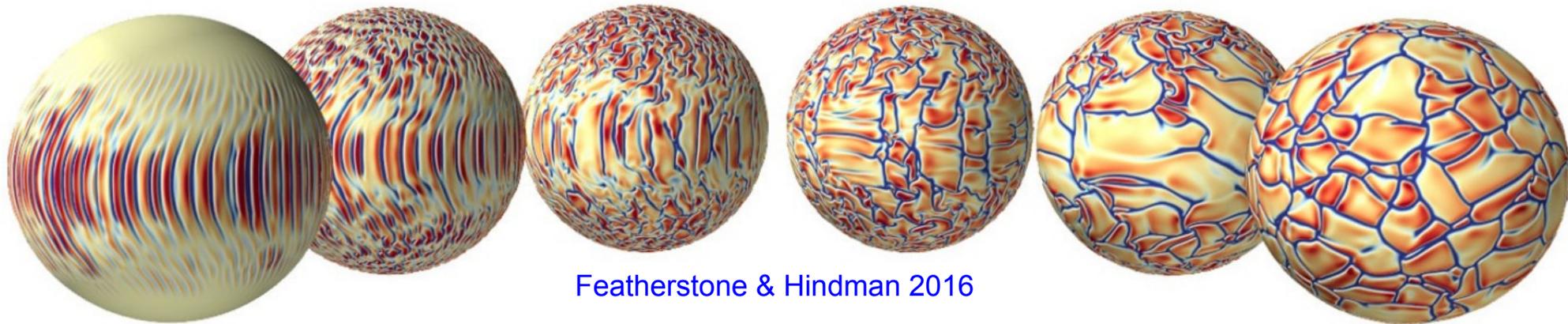


# Cycles grow longer and weaker in old stars



- Stalled rotation coincides with longer activity cycles and weaker variability
- Same pattern observed in hotter and cooler stars at same Rossby number
- Prediction: younger stars should have strong dipole while older stars will not

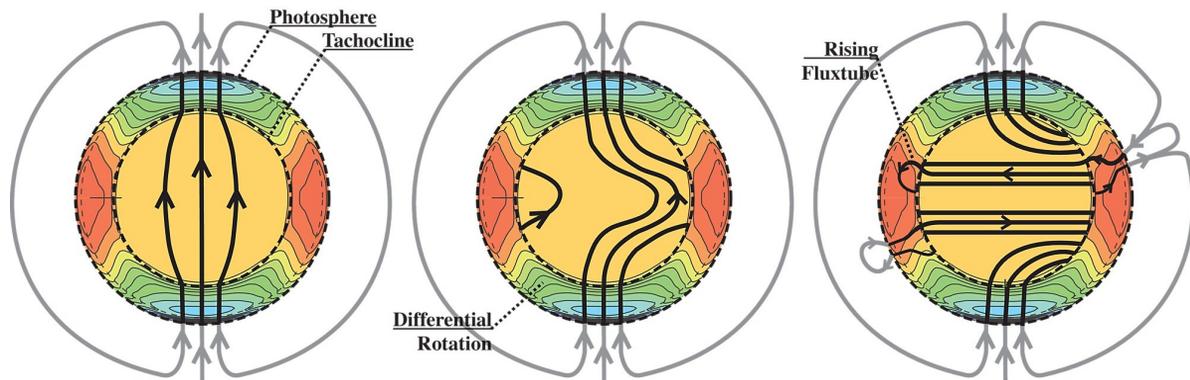
# 1. slow rotation becomes non-differential



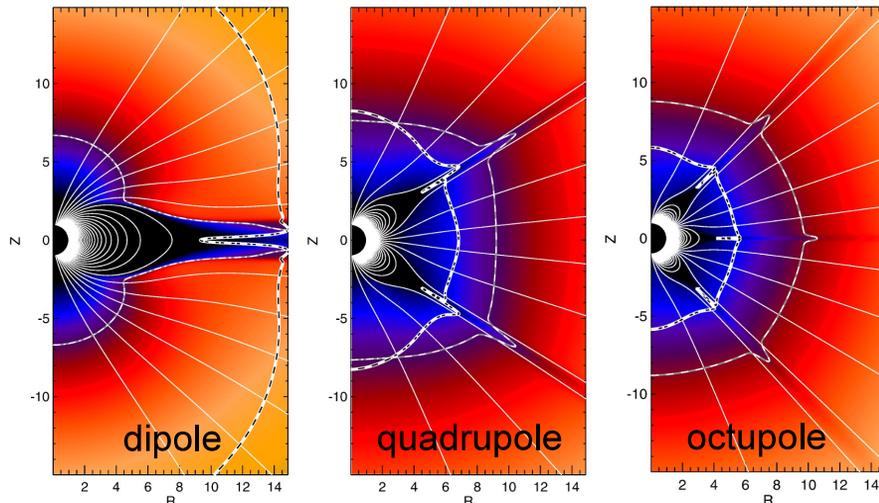
Featherstone & Hindman 2016

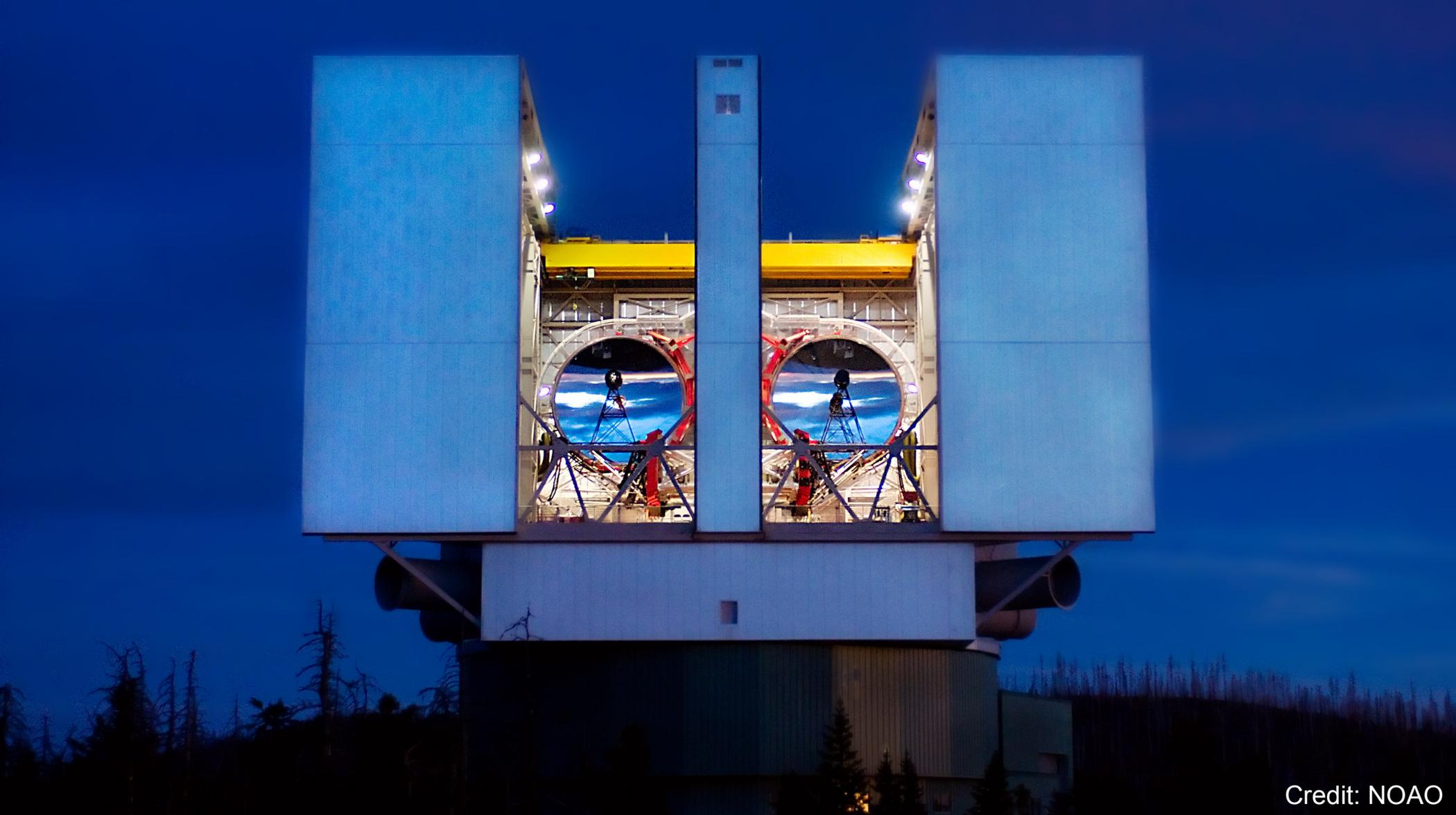
# 2. loss of shear disrupts $\alpha$ - $\Omega$ dynamo

# 3. decaying dipole stalls braking

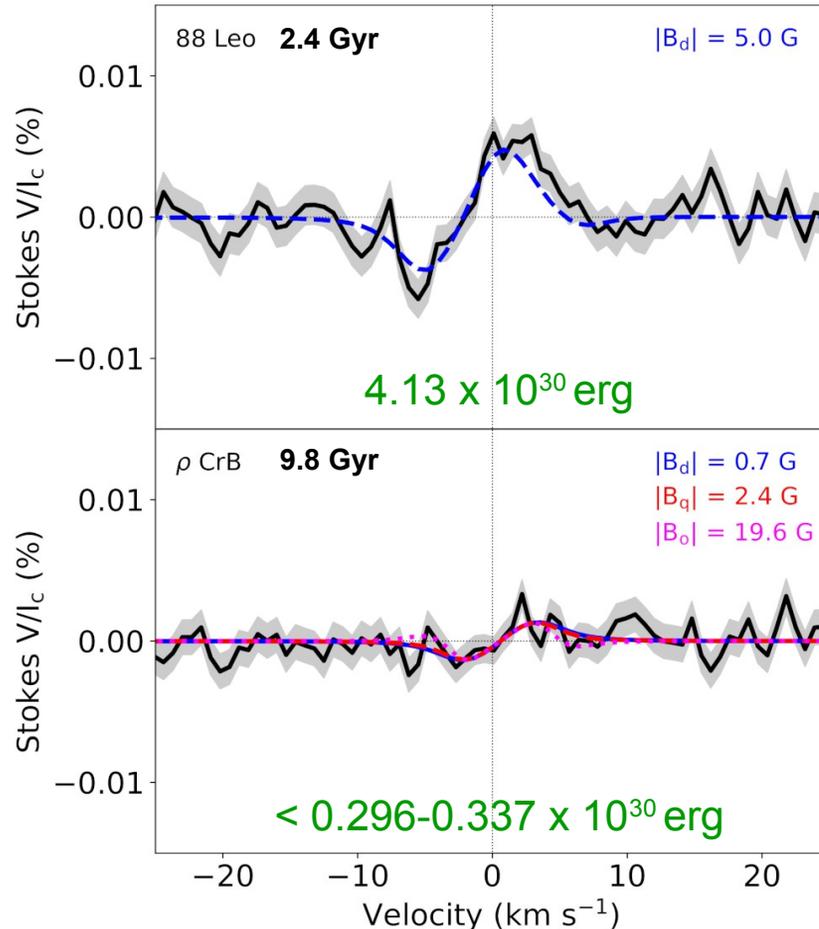


$\Omega$  effect (poloidal  $\rightarrow$  toroidal field)





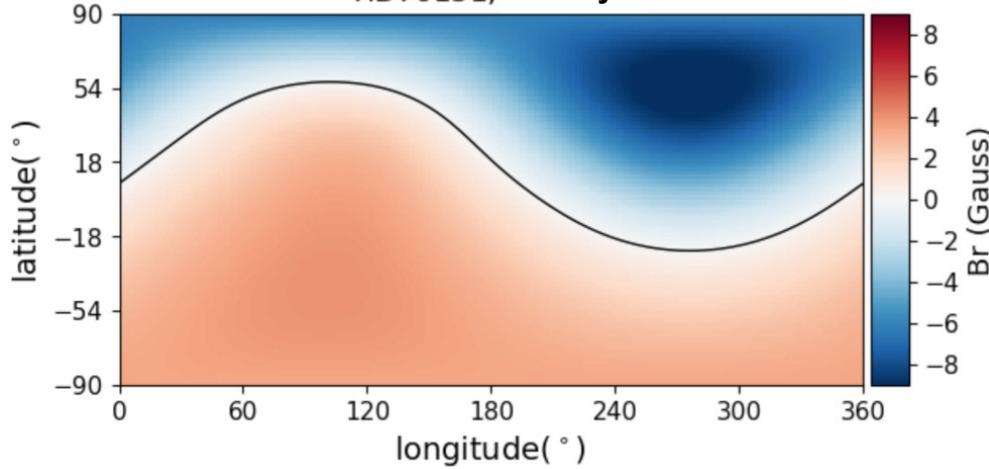
# Evolutionary sequence: F-type stars



- 88 Leo: detection of large-scale field, modeled by dipole with  $B_d = 5.0$  G
- $\rho$  CrB: upper limits on field strength suggest a torque  $< 8\%$  of 88 Leo
- Dominated by changes in field morphology, but ZDI needed for confirmation

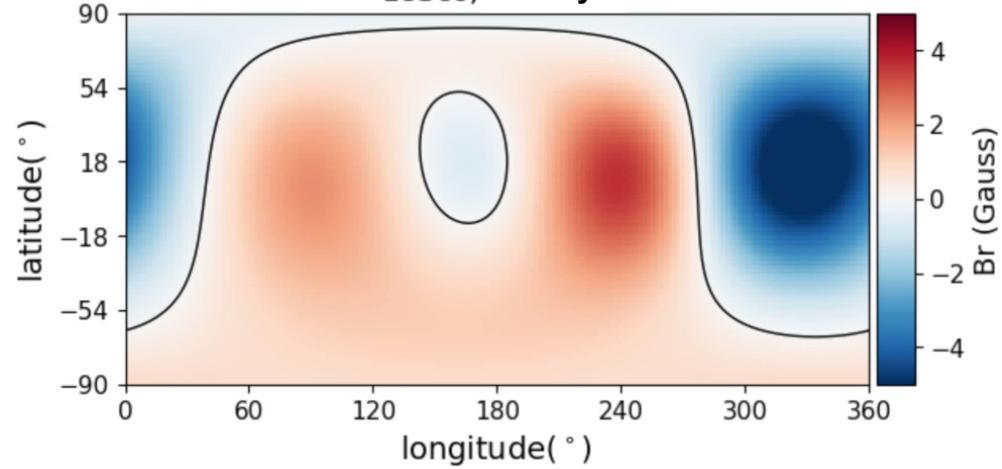
# Evolutionary sequence: solar analogs

HD76151, 2.6 Gyr



[ $B_d=5.13$  G,  $B_q=2.88$  G,  $B_o=1.34$  G]  $4.77 \times 10^{30}$  erg

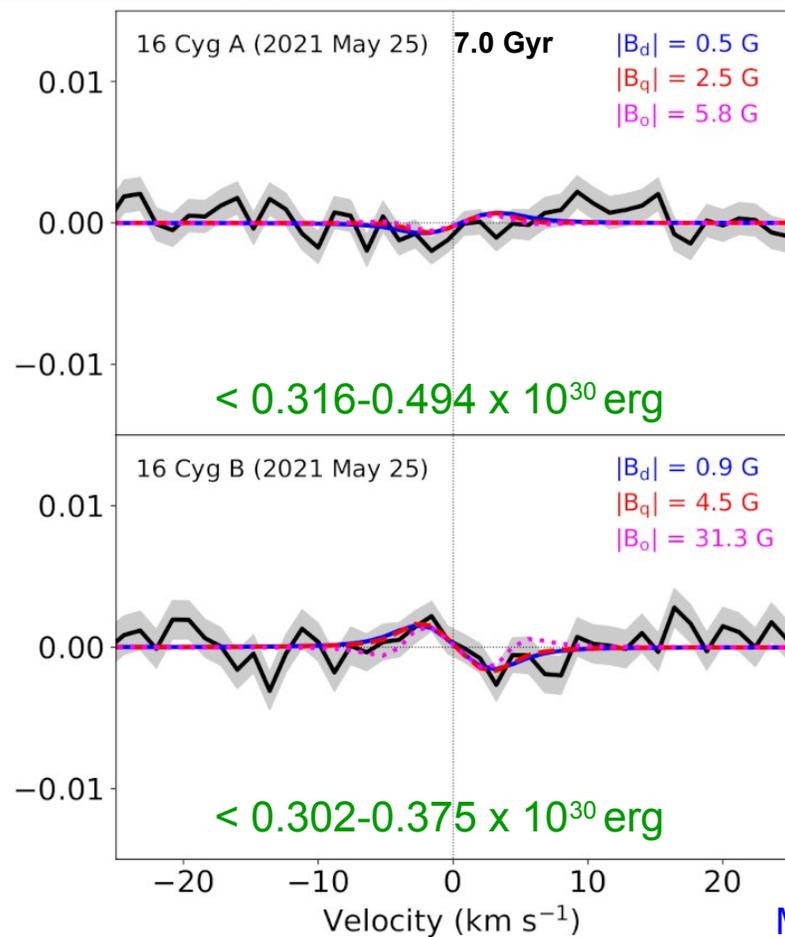
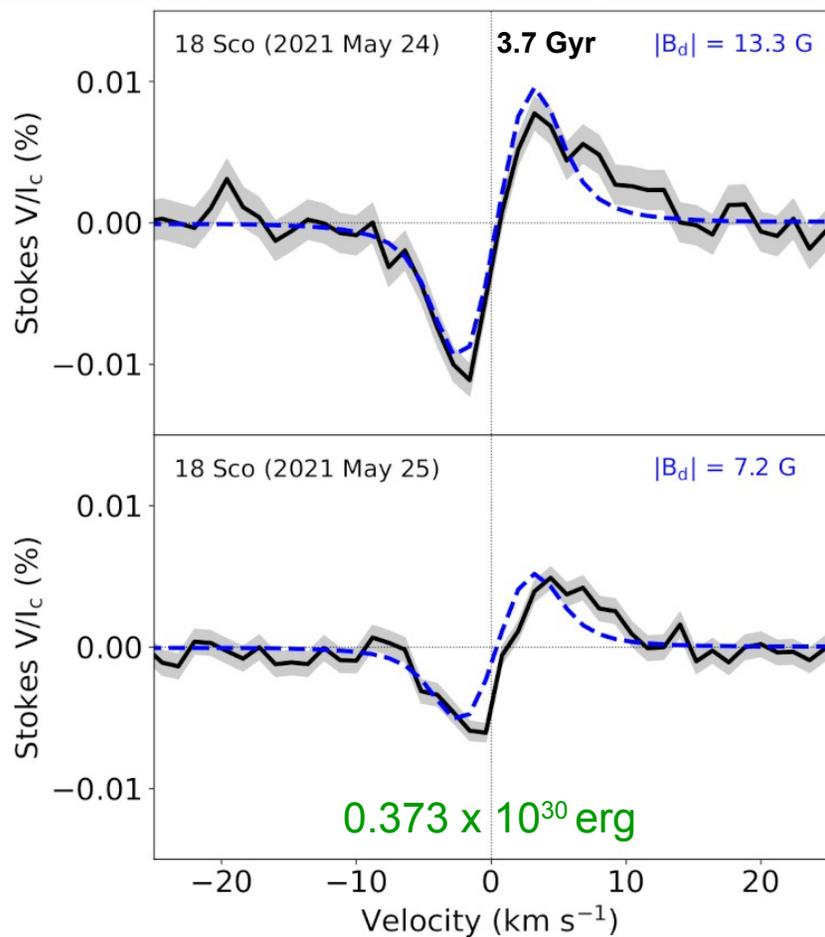
18Sco, 3.7 Gyr



[ $B_d=1.34$  G,  $B_q=2.01$  G,  $B_o=0.86$  G]  $0.373 \times 10^{30}$  erg

- Between 2.6-3.7 Gyr, braking torque decreases by an order of magnitude
- Dominated by changes in mass-loss rate and field strength & morphology

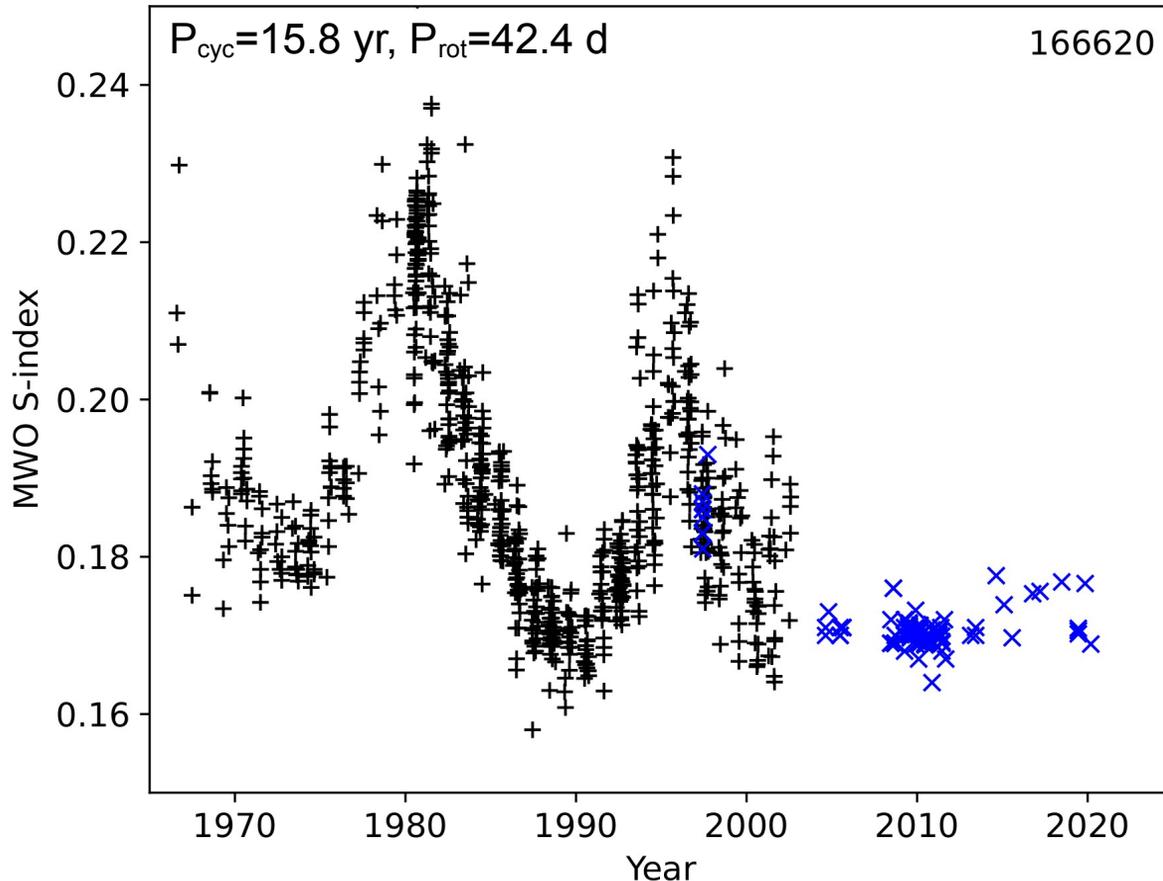
# Evolutionary sequence: old solar analogs



# Summary of conclusions

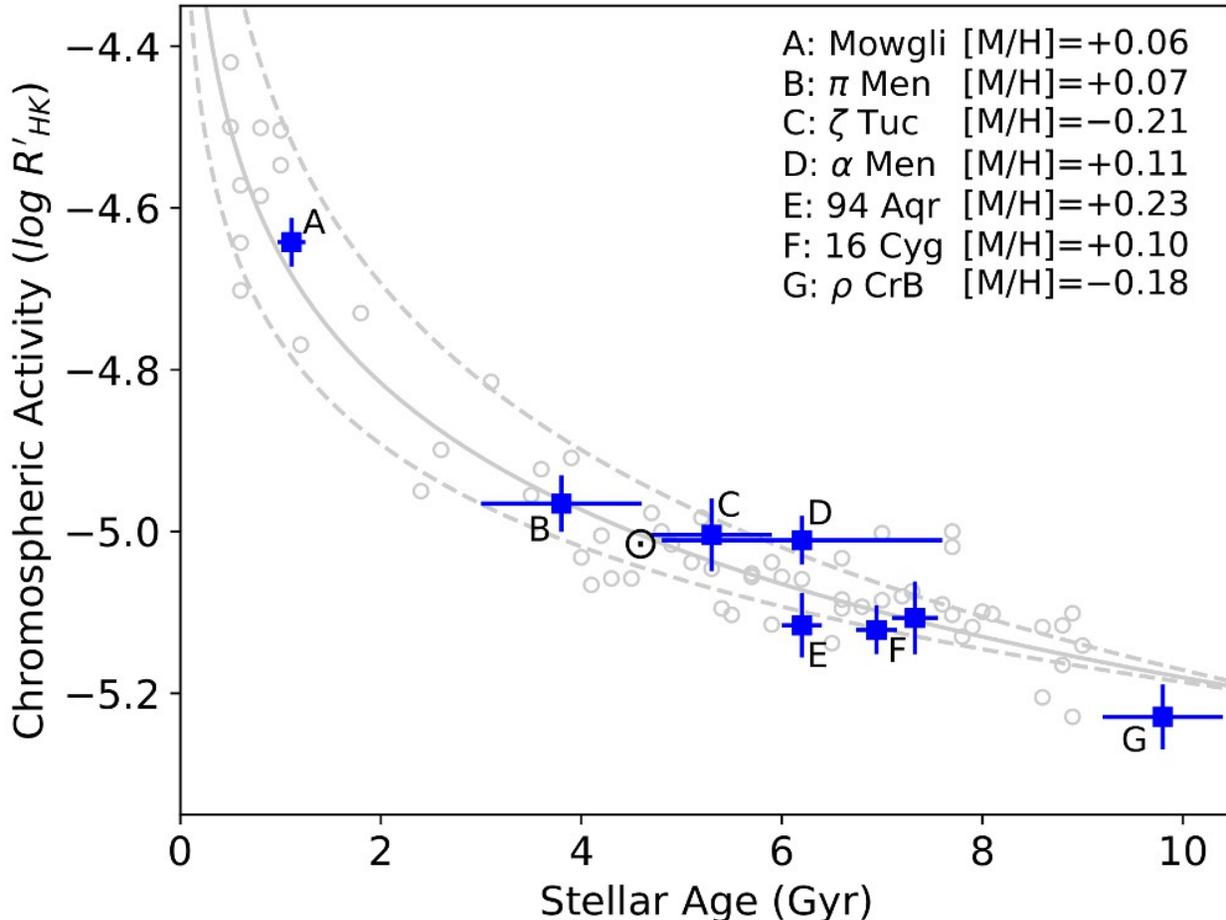
- Weakened magnetic braking beyond middle-age as the global field shifts toward smaller spatial scales
- F- and G-type stars make this transition at rotation periods between  $\sim 15$ -25 days, enabling direct tests
- LBT spectropolarimetry supports a diminishing dipole field in sequences of stars that span the transition
- Braking torque changes due to mass-loss rate (early) and field strength & morphology (dominates later)

# HD 166620: grand minimum



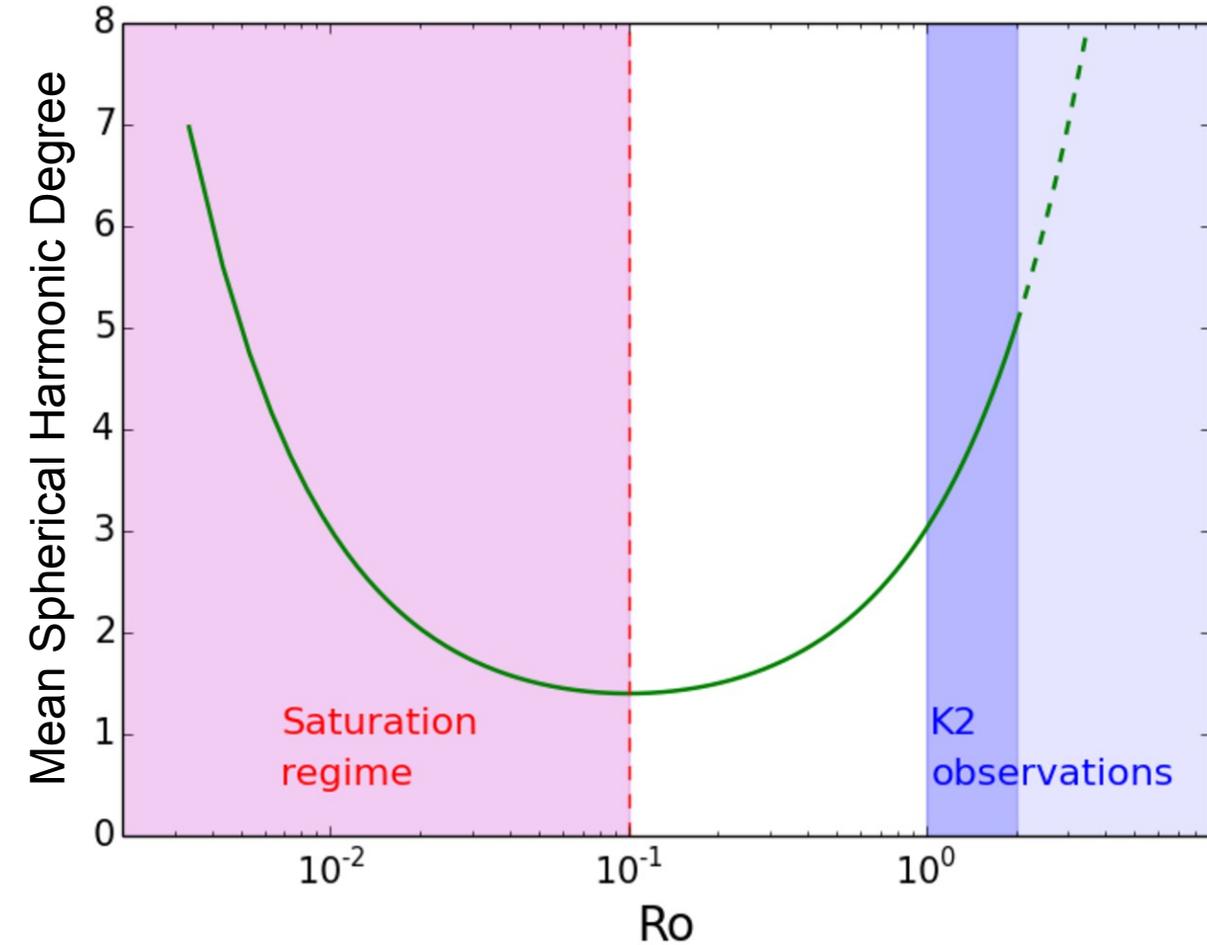
- Showed a clear Sun-like activity cycle during the Mount Wilson survey
- Keck data are consistent in the late-90s, constant activity level after 2003
- Position of cycle above lower sequence is outlier similar to the solar cycle

# Activity level evolves continuously with age



- Activity of solar analogs and asteroseismic targets decline continuously
- Solar dipole field is  $\sim 1$  G while unstructured quiet Sun has  $\langle B \rangle \sim 170$  G
- Disruption of large-scale organization is irrelevant to integrated activity level

# Evolution of magnetic complexity



- **Saturation regime:** range of rotation rates at constant activity level
- **Skumanich regime:** rotation rate and activity level decline together
- **Decoupled regime:** activity level evolves at constant rotation rate